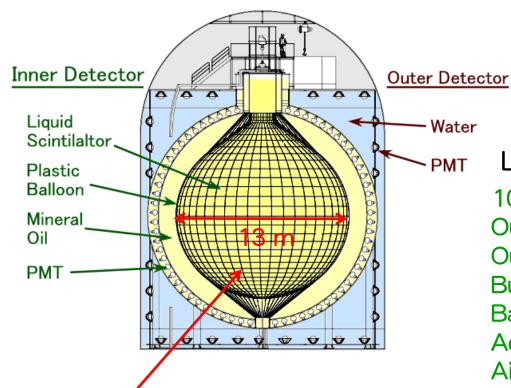
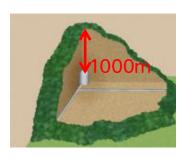
## KamLAND, KamLAND-Zen and Beyond

- 1. Double Beta Decay with KamLAND-Zen results, prospects and dreams
- 2. Evolution of KamLAND Scintillation Detector past 15 years and next 20 years

Sanshiro Enomoto University of Washington

### KamLAND Detector: Shielded Calorimeter





#### Layered Shields

1000 m underground for muon reduction Outer detector for active veto Outer detector for neutron shield Buffer oil for neutron/gamma shield Balloon with EVOH for Radon shield Acrylic wall for Radon shield Air tightness for Radon/Kr shield

Active Volume

Large: 1000 ton, 13-m diameter

Clean: U:  $3.5 \times 10^{-18}$  g/g, Th:  $5.2 \times 10^{-17}$  g/g

Scintillator: 8000 photons/MeV

Viewed by PMTs:  $1325 \times 17$ -inch +  $554 \times 20$ -inch

Originally constructed to detect antineutrinos from surrounding reactors "Evidence for Reactor Antineutrino Disappearance", PRL 90 021802 (2003)

### KamLAND as Active Low-BG Playground

#### Double Beta Decay

CdWO<sub>4</sub> Crystal

small crystal tested with KamLAND

CaF<sub>2</sub> Crystal

joint with CANDLES

Xe Dissolved Scintillator

KamLAND-Zen ...



#### Sterile Neutrinos

Ce Source (CeLAND)

- 75 kCi <sup>144</sup>Ce
- place center or outer vessel

#### Neutrino Beam (IsoDAR)

- use DAE δ ALUS injector
- 60 MeV proton beam to Be
- neutrons from Be
- <sup>7</sup>Li(n, γ )<sup>8</sup>Li
- 8Li beta decays: 10<sup>22</sup> / yr

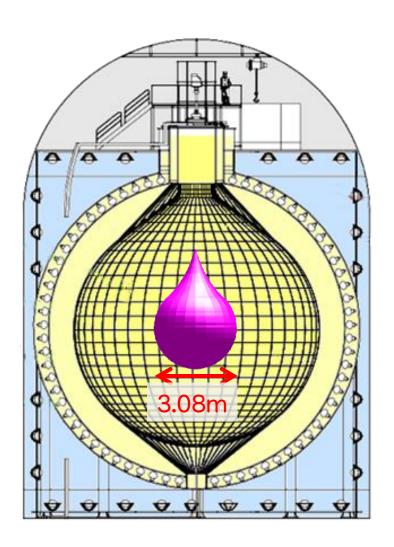
#### Dark Matter

Nal Crystal (KamLAND-PICO)

- verify DAMA/LIBRA results
- ~1000 kg Nal crystal
- U 1.1 ppt, Th 0.6 ppt crystal already developed

### KamLAND-Zen Project

Searches for  $0\nu\beta\beta$  with 380 kg of 91.7% Enriched <sup>136</sup>Xe (Q=2.458 MeV)

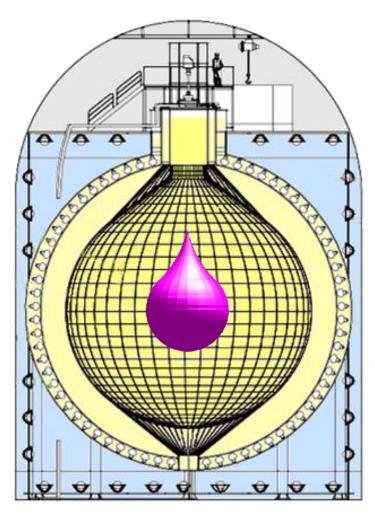


- √ 3 wt% Xe in liquid scintillator
- ✓ contained in 1.54m radius inner balloon
- √ balloon replaceable, Xe extractable
- ✓ make use of KamLAND facilities.
  - purification system
  - PMT, ATWD digitizer, analysis tools, ...
- ✓ can live with KamLAND (reactor & geo)
- √ No BG particle escape
  - no 2.6 MeV peak from <sup>208</sup>Tl
  - Bi-Po cascade decay tagging

#### **Drawbacks:**

- poor energy resolution (~4% @ 2.6 MeV)
- no  $\beta/\gamma$  discrimination

### KamLAND-Zen: Key Developments



#### Xe-loaded Liquid Scintillator

- Density balance
- Light-yield matching to outer LS
  - Xe reduces light-yield by ~10%
  - increase PPO
  - density ...

#### Inner Balloon

Not commercially available, developed for ZEN

- Clean: U/Th/K ~10<sup>-13</sup> g/g
   -no filler, no lubricant, no glue
- Thin:  $25 \mu m$  for Bi-Po  $\alpha$ -tagging
- Welding by heat-only (no glue)
- Xe-gas tight (<220 g/yr)</li>
- Transparent: 99.4% @400nm

### Inner-Balloon Fabrication & Installation

Ultrasonic cleaning in a super clean room







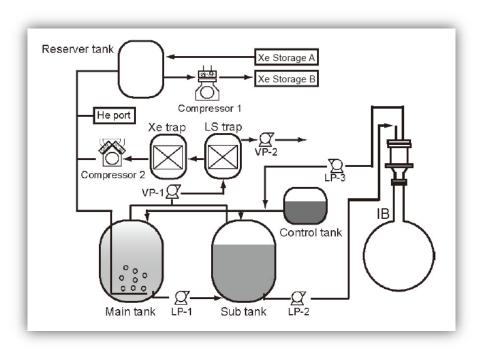








### Xe Handling Facility



- ✓ Xe Dissolving by bubbling
- ✓ Xe Extraction by degassing & bubbling with scintillator cold trap
- ✓ Precise Control & Monitor (0.005 ~ 0.01 %)
  by density measurement & mass balance
- ✓ 91.7% Enriched <sup>136</sup>Xe, 610 kg in hand







### Xe-LS Filling into Inner-Balloon

Inner balloon was initially filled with z [m] normal LS (no Xe) Then Xe-LS was fed from bottom 0.5 with slightly higher density (+0.035%) Initial Rn contamination makes -0.5newly filled Xe-LS very visible enrXe 330 kg filled 1.5 2.5  $x^2 + y^2 [m^2]$ (enrichment 91.7%)

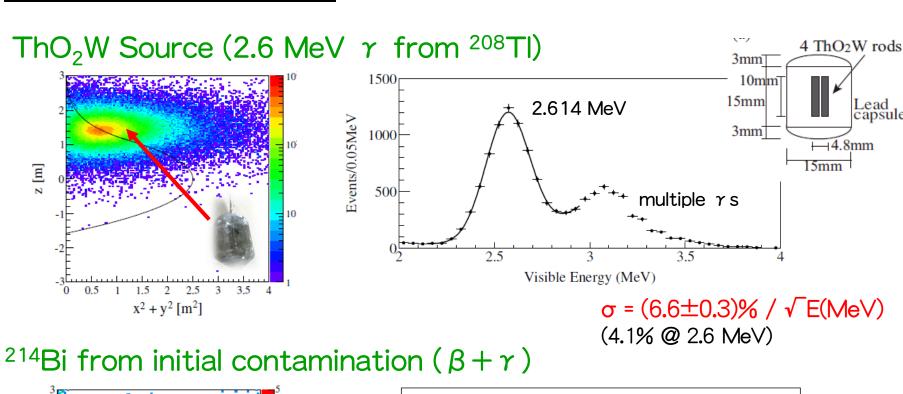
 $x^2 + y^2 [m^2]$ 

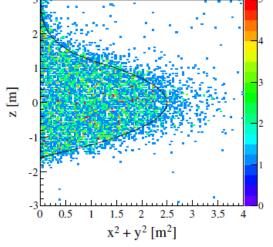
Start data taking on Sep 24th 2011 (only two years from the first Zen funding)

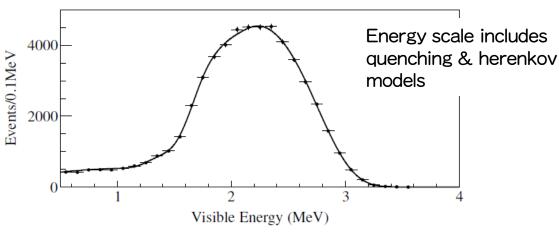
Mix the whole volume

by few more full circulation

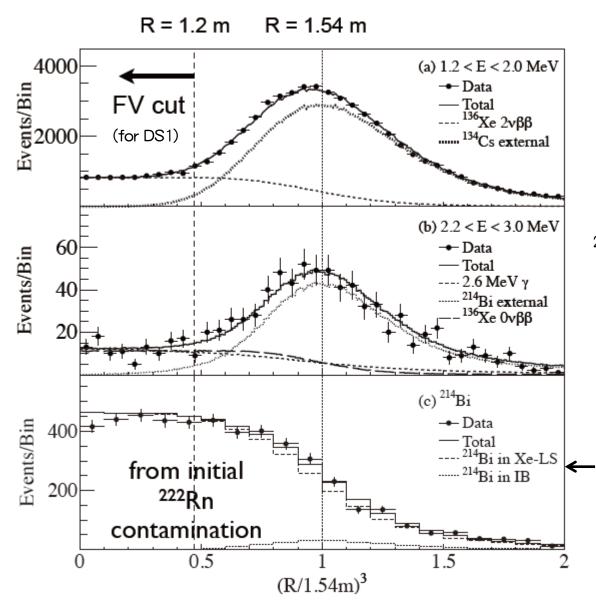
### **Energy Calibration**

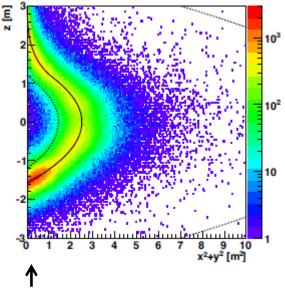






### Fiducial Volume





<sup>214</sup>Bi in balloon limits FV

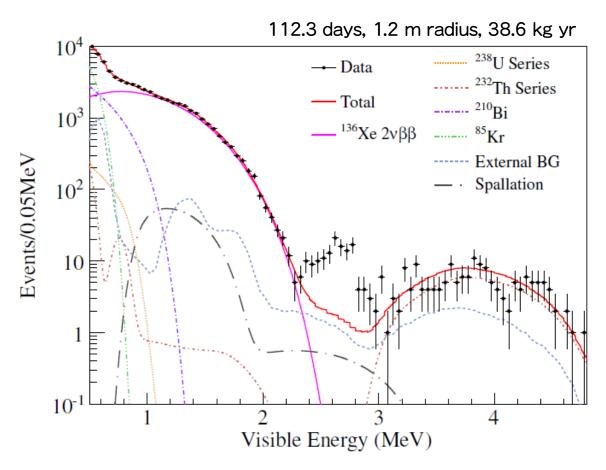
Vertex resolution:  $\sigma \sim 15 \text{ cm} / \sqrt{\text{E(MeV)}}$ 

FV ratio estimated with uniformly distributing initial <sup>214</sup>Bi

$$\Rightarrow$$
 3.9% / 4.1% FV error (DS1) (DS2)

All other syst. errors: <1% (136Xe amount, efficiency, etc)

### Energy Spectrum & 2νββ Half Life



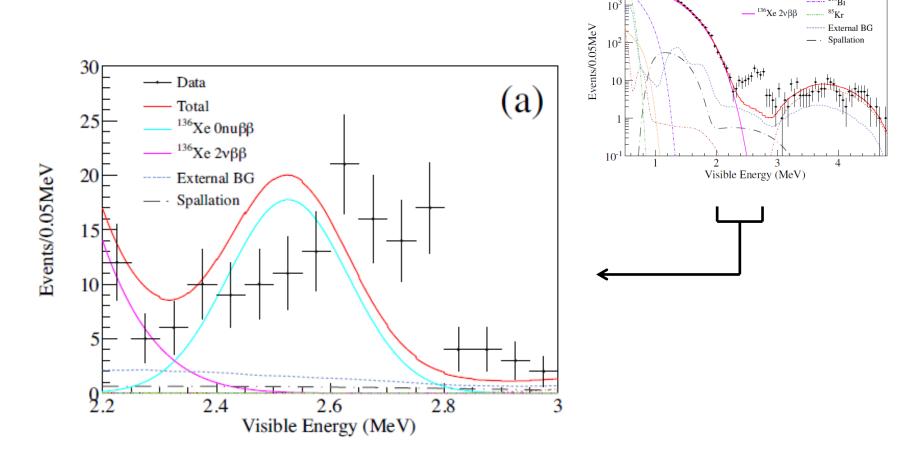
$$T_{1/2}^{2\nu}$$
 = 2.38 ± 0.02 (stat) ± 0.14 (syst) ×10<sup>25</sup> yr

Consistent with EXO-200

Phys.Rev.C 85 045504 (2012)

### The Peak is not 0νββ

(and not <sup>208</sup>TI)



 $10^{4}$ 

238U Series

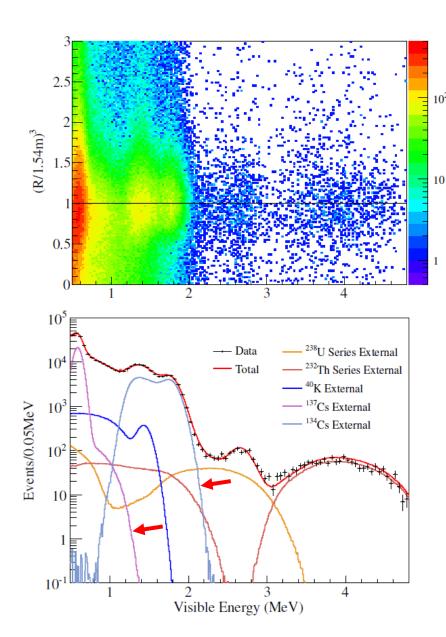
··· <sup>232</sup>Th Series

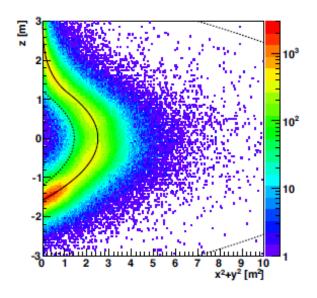
→ Data

— Total

 $0\nu$ ββ Spectrum shape is inconsistent at >8 σ

## Balloon is Contaminated with Fukushima Fallout



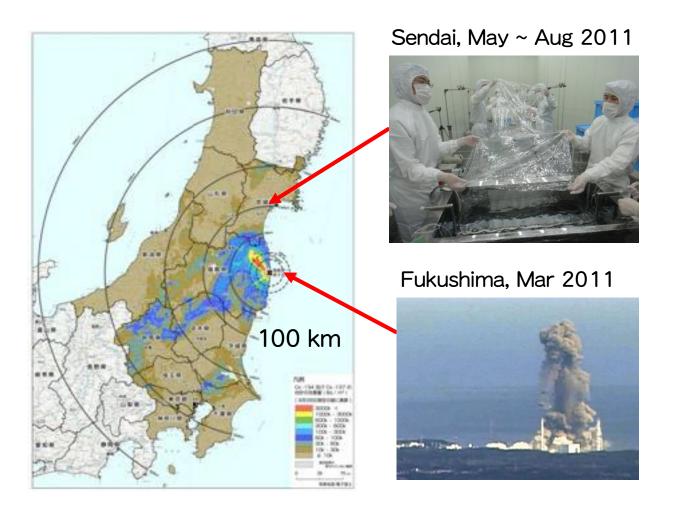


<sup>137</sup>Cs and <sup>134</sup>Cs dominate the balloon activity

 $^{134}\text{Cs}$  /  $^{137}\text{Cs}$  ~ 0.8, consistent with fallout composition

Cs is not soluble in scintillator

### Balloon is Contaminated with Fukushima Fallout

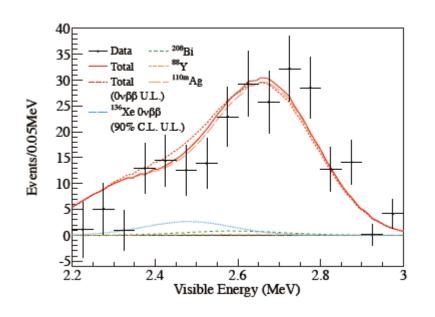


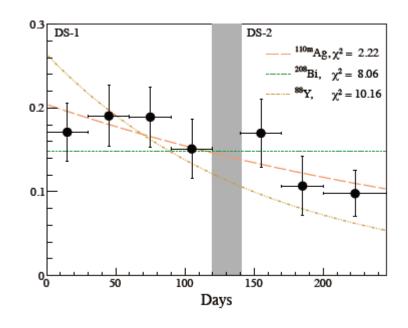
Fallouts detected in Sendai Soil:

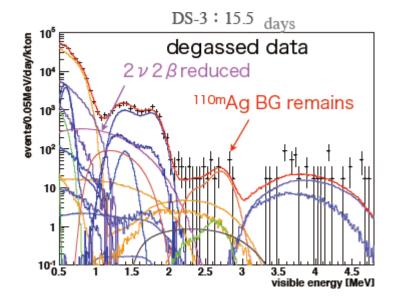
134Cs, 137Cs, 110mAg, 129mTe, 95Nb, 90Y, 89Sr

→ all included in spectrum fit

## ~1000 atoms of 110mAg in Volume



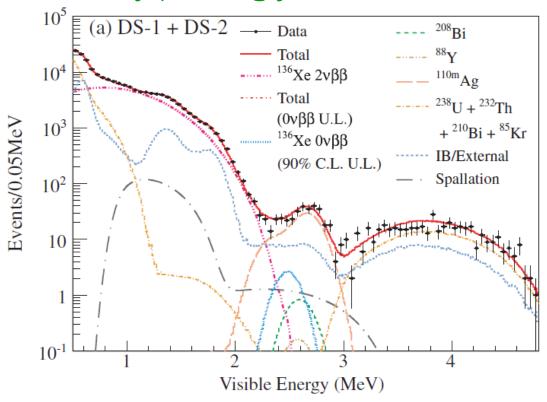




- ✓ Spectrum most consistent with <sup>110m</sup>Ag (<sup>208</sup>Bi, <sup>88</sup>Y do not fit well, but not excluded)
- ✓ Decay-time consistent with <sup>110m</sup>Ag (τ<sub>1/2</sub>: 250 days)
- ✓ Peak remains after Xe extraction → not from Xe - cosmic- $\mu$  interaction

## <u>0νββ Limit</u> PRL 110, 062502 (2013)

#### 213.4 days, 98.5 kg yr

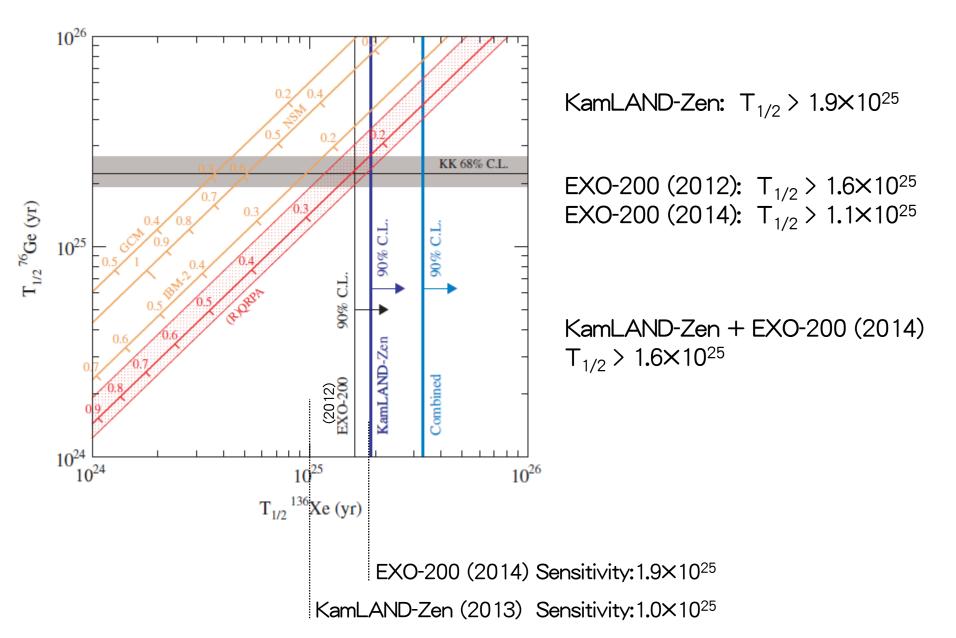


Likelihood fit to binned spectrum  $0\nu\beta\beta$ : < 0.16 events/kg/yr

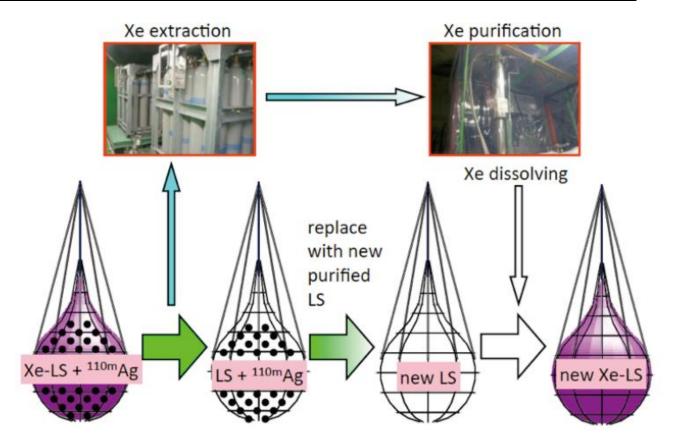
$$T_{1/2}^{0\nu}$$
 > 1.9×10<sup>25</sup> yr (90% CL)  $\langle m_{\beta\beta} \rangle$  < 120 ~ 250 meV

(sensitivity:  $1.0 \times 10^{25}$  yr; we are top 12% lucky)

## Combined with EXO-200, Comparison with Ge



## Ongoing: 110mAg Removal by Distillation

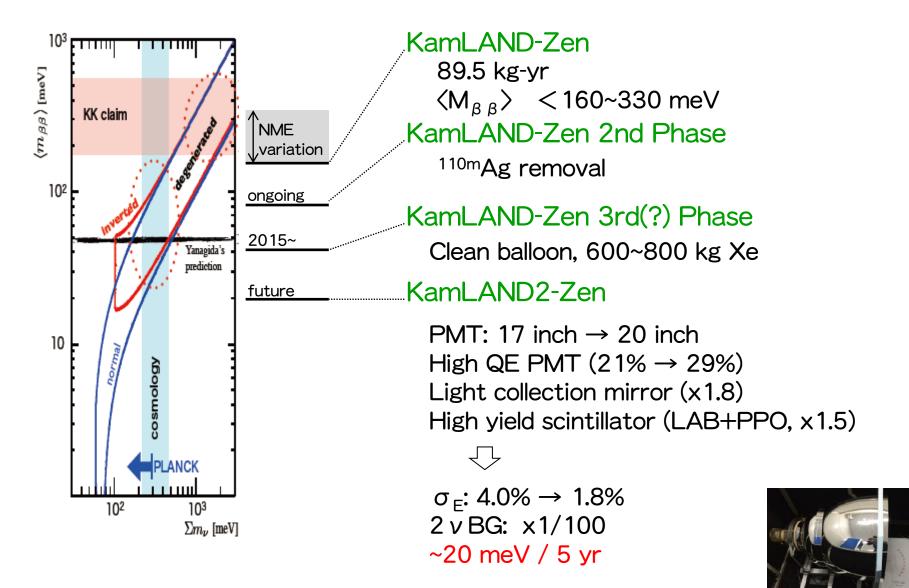


Completed in Nov 2013 >90% of 110mAg removed

#### KamLAND-Zen 2nd Phase

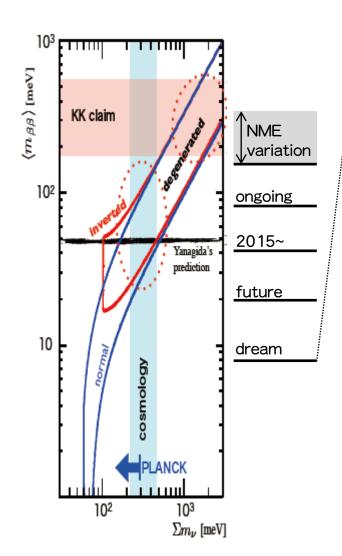
Dec 2013 ~, enrXe 380 kg (+60 kg to phase 1)

### Zen Future: Less Background, More Resolution



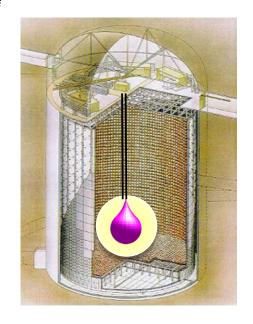
mirror prototype

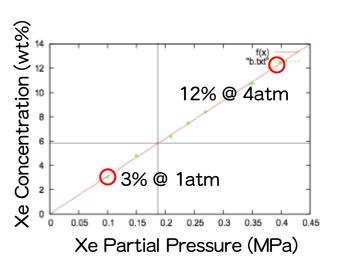
### Zen Dream: More Target Density



Super-KamLAND-Zen

Pressurize the Zen Balloon



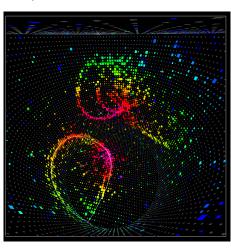


× 4 more Xe concentration at 30 m deep

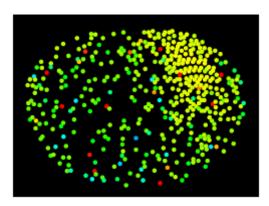
### Evolution of KamLAND Scintillation Detector

#### Water Cherenkov Detector vs Scintillation Detector

#### Super-Kamiokande



#### **KamLAND**



#### Advantages

Clean (U & Th are ionic)

High light yield (x ~100 of Water Cherenkov)

→ Low energy threshold (~300 eV, limited by <sup>14</sup>C)

Delayed coincidence  $\rightarrow$  Selective sensitivity to  $\overline{V}_e$ 

#### **Drawbacks**

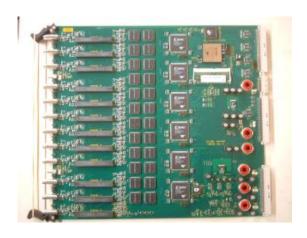
Costly + Layered Shields → Small

No directional sensitivity

Signal pile-up

Mess after cosmic muons (more p.e. + low thresh)

## Pile-up Free Electronics (LBL, 2001)



#### Waveform Recording with ATWD

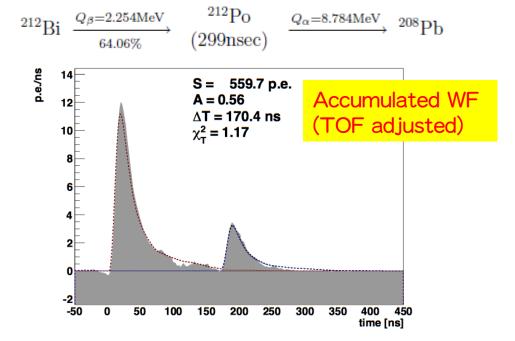
(ATWD: Analog Transient Waveform Digitizer)

~1.5 ns interval sampling for all channels!!

Two 4-ch ATWDs per PMT

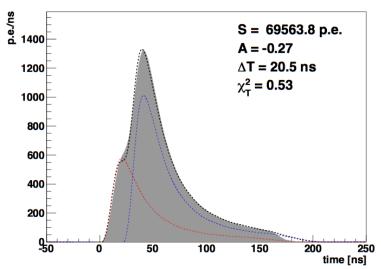
- 4-ch for dynamic range (x20, x4 and x0.5 plus CLK)
- two ATWDs for cascading events

#### Bi-Po Pile-up Waveform (Data)



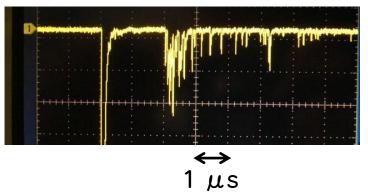
#### Proton-Decay Waveform (MC)

$$p \rightarrow K^+ \overline{\nu}_\mu; K^+ \rightarrow \mu^+ \nu_\mu; \mu^+ \rightarrow e^+ \overline{\nu}_\mu \nu_e$$

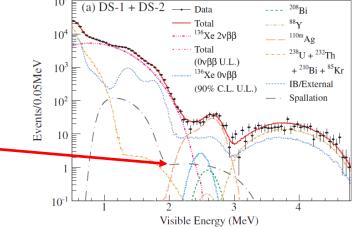


### Dead-time Free Electronics (Tohoku, 2009)

#### Mess After Cosmic Muons



- after-pulses: up-to ~100k events in ~100  $\mu$ s
- neutrons: up-to ~100 events in ~ms



neutrons tells us location of  $^{10}$ C  $\rightarrow$  pin-point veto

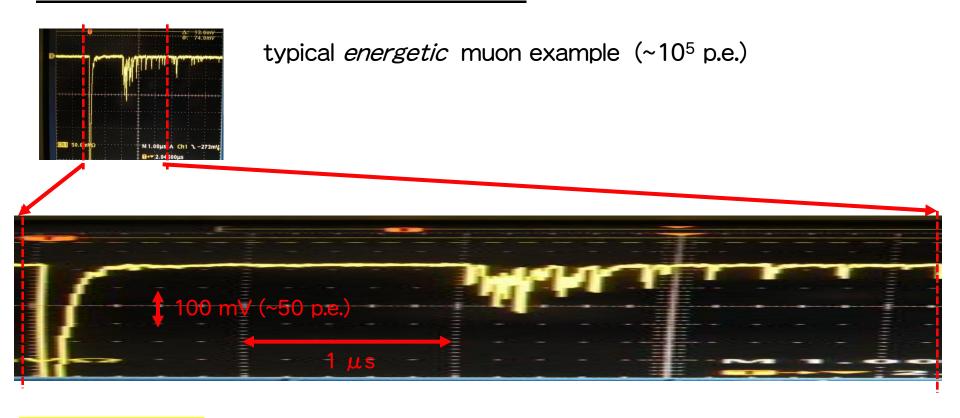
#### Free-Running FADC + Inline Digital Processing

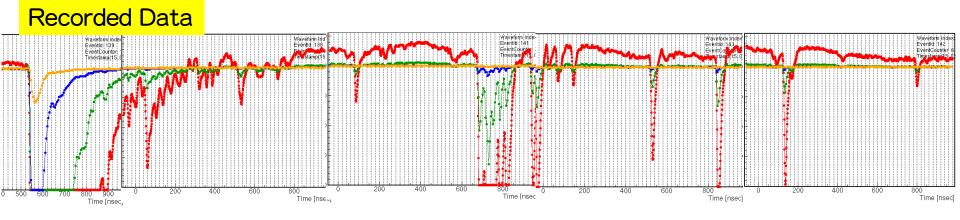


1 GHz FADC per ch10 μs ring buffer per ch>256 event buffer per ch

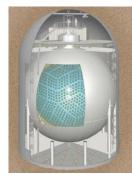
Very similar to the former design after 8 years!!

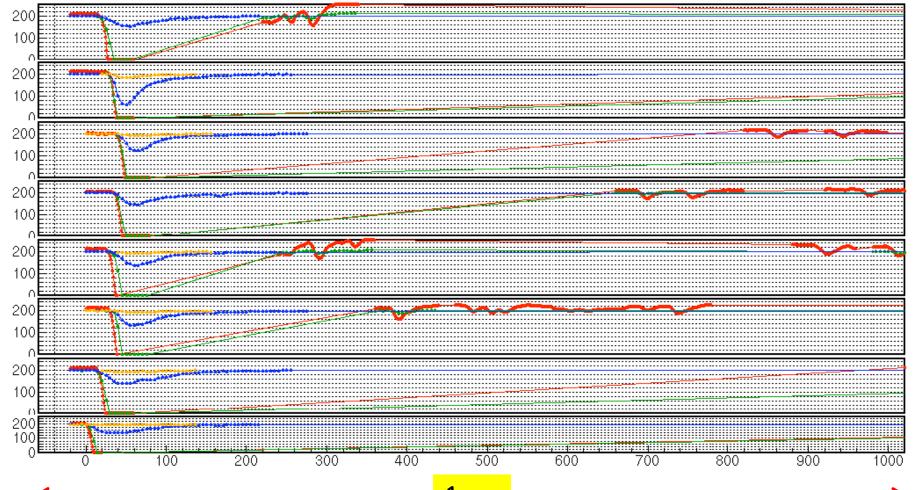
Inline data processing (zero suppression etc)
Onboard processor, serial readout option



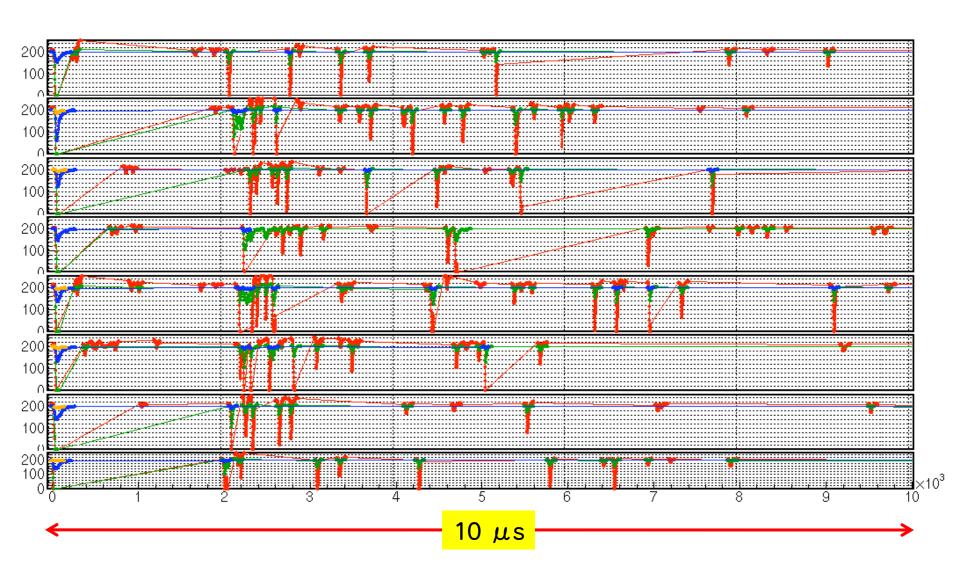


All 1325 channels, with zero-suppression (only 8 ch are shown below)

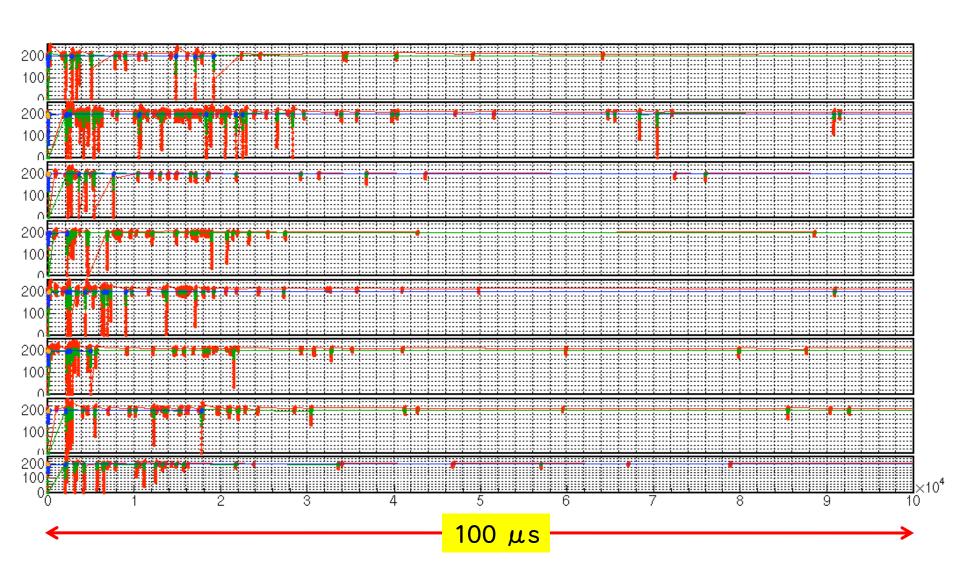




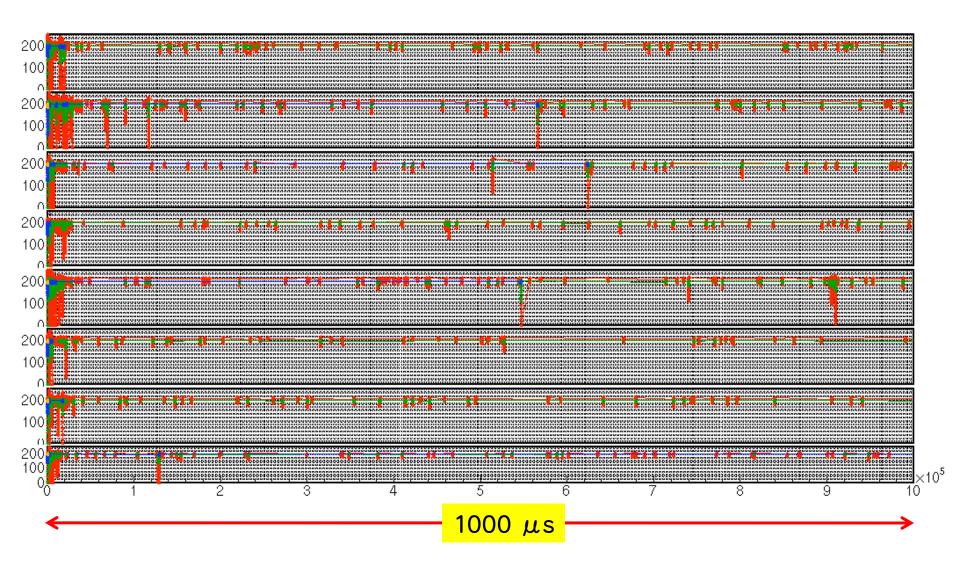
In a longer time range



### Even longer time range

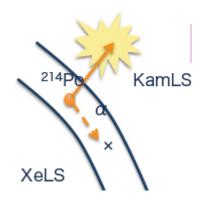


#### And more!!



### <u>Dead-Layer Free Scintillating Balloon (R&D)</u>

# Imperfect Bi-Po tagging due to balloon *dead-layer*



#### PEN was found to be scintillating material (2011)



Material	Polyethylene naphthalate	Organic scintillator (ref. [14])
Supplier	Teijin Chemicals	Saint-Gobain
Base	$(C_{14}H_{10}O_4)_n$	$(C_9H_{10})_n$
Density	$(C_{14}H_{10}O_4)_n$ $1.33 \text{ g/cm}^3$	$1.03{\rm g}/{\rm cm}^3$
Refractive index	1.65	1.58
Light output	$\sim 10500~\rm photon/MeV$	$10000 \mathrm{\ photon/MeV}$
Wavelength max. emission	$425\mathrm{nm}$	$425\mathrm{nm}$

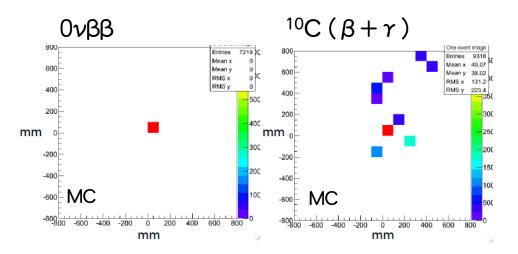


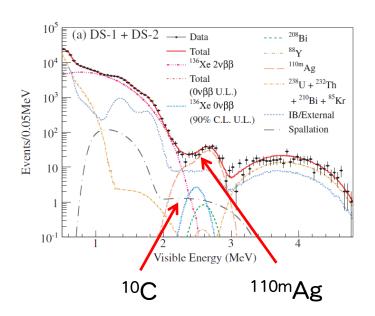
#### First look at 50 $\mu$ m-thick PEN film (Tohoku)

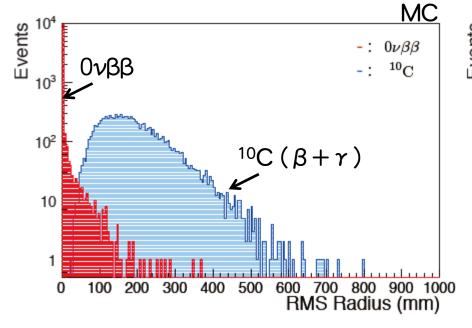
- √ Scintillates!!
- ✓ Clean enough, Weld-able, chemically compatible with LS Detail studies ongoing

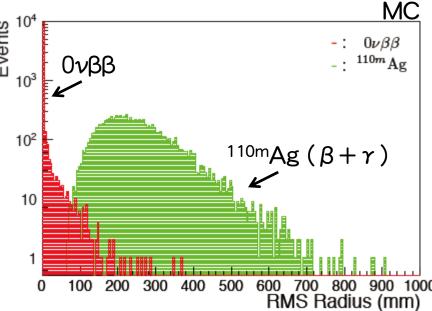
## Scintillation Imaging (R&D)

#### Particle Identification using Event Size



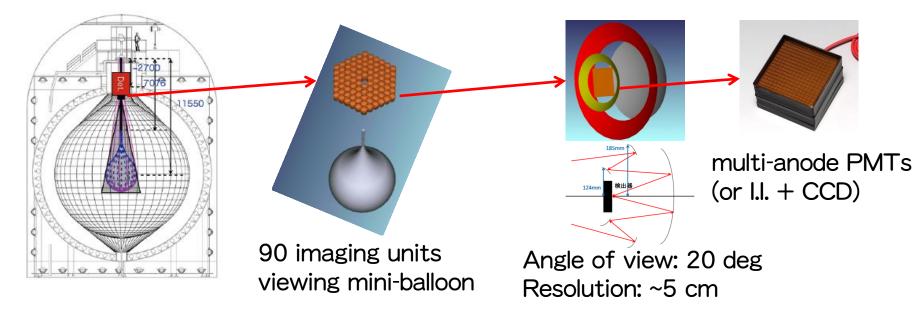




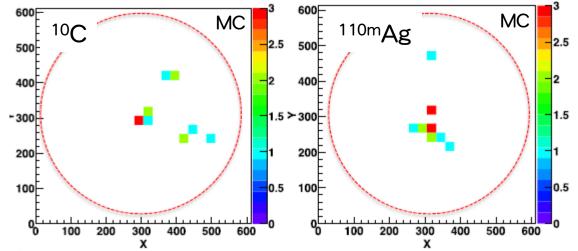


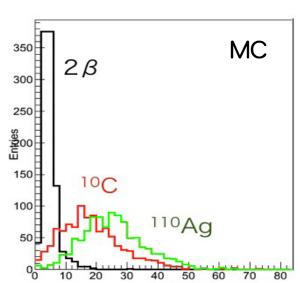
## Scintillation Imaging (R&D)

Optics and Readout Design (Tohoku)



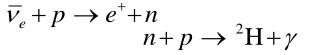


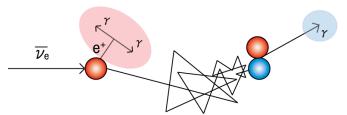




## Directional Sensitivity (R&D)

#### Scintillation Imaging for Anti-Neutrino Events





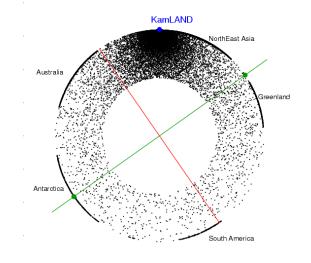
Neutrons remember neutrino direction but forget it soon

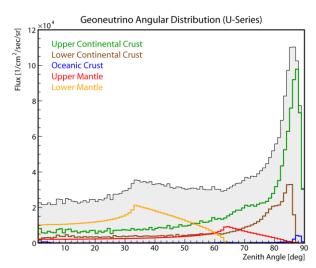
For Directionality,

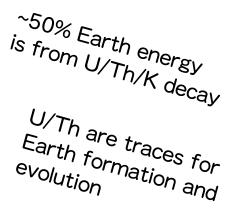
- ✓ Reduce neutron diffusion
- ✓ Eliminate gamma emission

### Detecting Antineutrino Direction is a Revolution in Geophysics

3D Mapping of U & Th distributions in Earth interior (with multi-site observation, or mobile detectors)







### Directional Sensitivity (R&D)

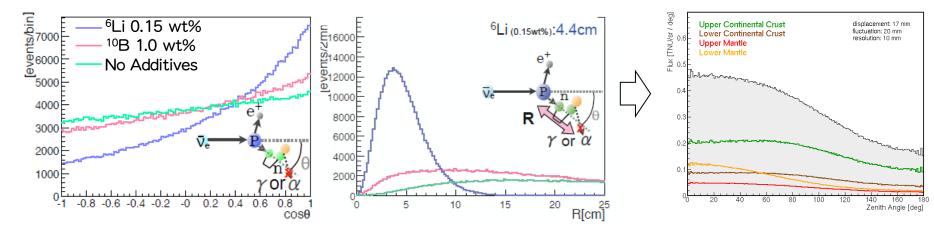
#### <sup>6</sup>Li-loaded Liquid Scintillator

$$\overline{v}_e + p \rightarrow e^+ + n$$

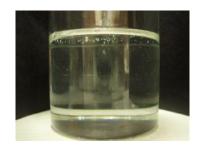
$$n + {}^6\text{Li} \rightarrow {}^3\text{H} + \alpha$$

- ✓ Large capture cross-section 940 barn (H: 0.3 barn)
- ✓  $\alpha$ -emission, no  $\gamma$  (Gd and <sup>10</sup>B both emit  $\gamma$ )

#### MC Simulation (0.15 wt% <sup>6</sup>Li)



#### Li-loaded Liquid Scintillator Development (Tohoku)

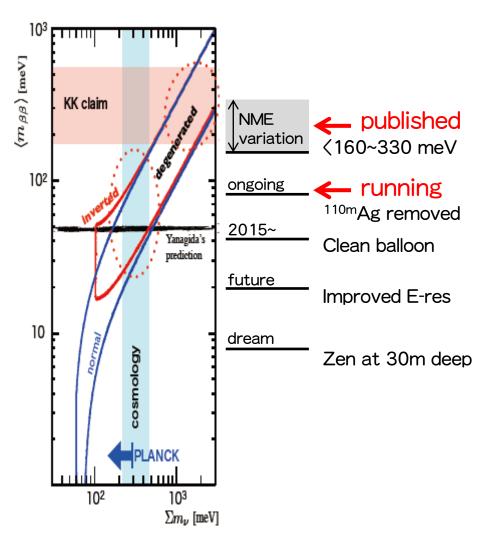


 $PC + PPC + POE + LiBr + H_2O$  (POE: surfactant)

0.8 wt% Li (6Li abundance 7.59%, enrichment is possible) Transparency 65cm@400nm, light yield 45% of KamLAND LS

### <u>Summary</u>

#### KamLAND-Zen 0νββ



# KamLAND Scintillation Detector Evolution

#### 2002~

- Large Volume, Ultra-Pure
- Pile-up Free Readout (ATWD)

#### 2010~

 Dead-time Free Readout for <sup>10</sup>C tagging etc

#### R&D

 Dead-space Free Container with scintillating balloon

#### R&D

• Scintillation Imaging for  $e/\gamma$  discrimination etc

#### R&D

Directional Sensitivity
 for geoneutrinos, supernova, ···

### KamLAND Founders

VOLUME 90, NUMBER 2

PHYSICAL REVIEW LETTERS

week ending 17 JANUARY 2003

#### First Results from KamLAND: Evidence for Reactor Antineutrino Disappearance

K. Eguchi, <sup>1</sup> S. Enomoto, <sup>1</sup> K. Furuno, <sup>1</sup> J. Goldman, <sup>1</sup> H. Hanada, <sup>1</sup> H. Ikeda, <sup>1</sup> K. Ikeda, <sup>1</sup> K. Inoue, <sup>1</sup> K. Ishihara, <sup>1</sup> W. Itoh, <sup>1</sup> T. Iwamoto, <sup>1</sup> T. Kawashima, <sup>1</sup> H. Kinoshita, <sup>1</sup> Y. Kishimoto, <sup>1</sup> M. Koga, <sup>1</sup> Y. Koseki, <sup>1</sup> T. Maeda, <sup>1</sup> T. Mitsui, <sup>1</sup> M. Motoki, <sup>1</sup> K. Nakajima, <sup>1</sup> M. Nakajima, <sup>1</sup> T. Nakajima, <sup>1</sup> H. Ogawa, <sup>1</sup> K. Owada, <sup>1</sup> T. Sakabe, <sup>1</sup> I. Shimizu, <sup>1</sup> J. Shirai, <sup>1</sup> F. Suekane, <sup>1</sup> A. Suzuki, <sup>1</sup> K. Tada, <sup>1</sup> O. Tajima, <sup>1</sup> T. Takayama, <sup>1</sup> K. Tamae, <sup>1</sup> H. Watanabe, <sup>1</sup> J. Busenitz, <sup>2</sup> Z. Djurcic, <sup>2</sup> K. McKinny, <sup>2</sup> D.-M. Mei, <sup>2</sup> A. Piepke, <sup>2</sup> E. Yakushev, <sup>2</sup> B. E. Berger, <sup>3</sup> Y. D. Chan, <sup>3</sup> M. P. Decowski, <sup>3</sup> D. A. Dwyer, <sup>3</sup> S. J. Freedman, <sup>3</sup> Y. Fu, <sup>3</sup> B. K. Fujikawa, <sup>3</sup> K. M. Heeger, <sup>3</sup> K. T. Lesko, <sup>3</sup> K.-B. Luk, <sup>3</sup> H. Murayama, <sup>3</sup> D. R. Nygren, <sup>3</sup> C. E. Okada, <sup>3</sup> A.W. P. Poon, <sup>3</sup> H. M. Steiner, <sup>3</sup> L. A. Winslow, <sup>3</sup> G. A. Horton-Smith, <sup>4</sup> R. D. McKeown, <sup>4</sup> J. Ritter, <sup>4</sup> B. Tipton, <sup>4</sup> P. Vogel, <sup>4</sup> C. E. Lane, <sup>5</sup> T. Miletic, <sup>5</sup> P.W. Gorham, <sup>6</sup> G. Guillian, <sup>6</sup> J. G. Learned, <sup>6</sup> J. Maricic, <sup>6</sup> S. Matsuno, <sup>6</sup> S. Pakvasa, <sup>6</sup> S. Dazeley, <sup>7</sup> S. Hatakeyama, <sup>7</sup> M. Murakami, <sup>7</sup> R. C. Svoboda, <sup>7</sup> B. D. Dieterle, <sup>8</sup> M. DiMauro, <sup>8</sup> J. Detwiler, <sup>9</sup> G. Gratta, <sup>9</sup> K. Ishii, <sup>9</sup> N. Tolich, <sup>9</sup> Y. Uchida, <sup>9</sup> M. Batygov, <sup>10</sup> W. Bugg, <sup>10</sup> H. Cohn, <sup>10</sup> Y. Efremenko, <sup>10</sup> Y. Kamyshkov, <sup>10</sup> A. Kozlov, <sup>10</sup> Y. Nakamura, <sup>10</sup> L. De Braeckeleer, <sup>11</sup> C. R. Gould, <sup>11</sup> H. J. Karwowski, <sup>11</sup> D. M. Markoff, <sup>11</sup> J. A. Messimore, <sup>11</sup> K. Nakamura, <sup>11</sup> R. M. Rohm, <sup>11</sup> W. Tornow, <sup>11</sup> A. R. Young, <sup>11</sup> and Y.-F. Wang<sup>12</sup>

(KamLAND Collaboration)